# IRS Observations of Water Vapor (and other gas phase molecules)

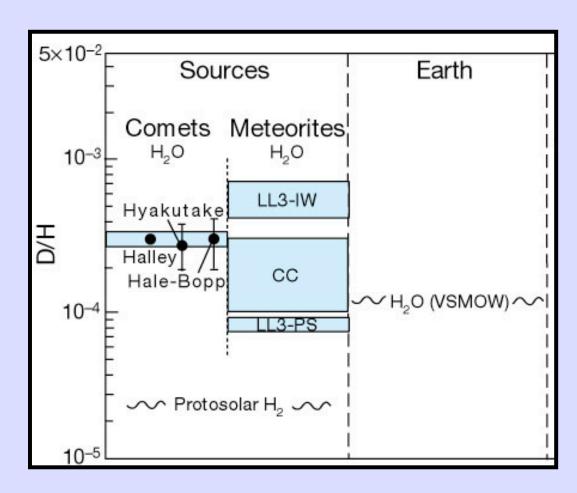
### Colette Salyk

Geoffrey Blake, John Carr, Ewine van Dishoeck, Fred Lahuis, Joan Najita, Klaus Pontoppidan, the c2d Legacy team

#### What is the source of Earth's water?

Primordial source?

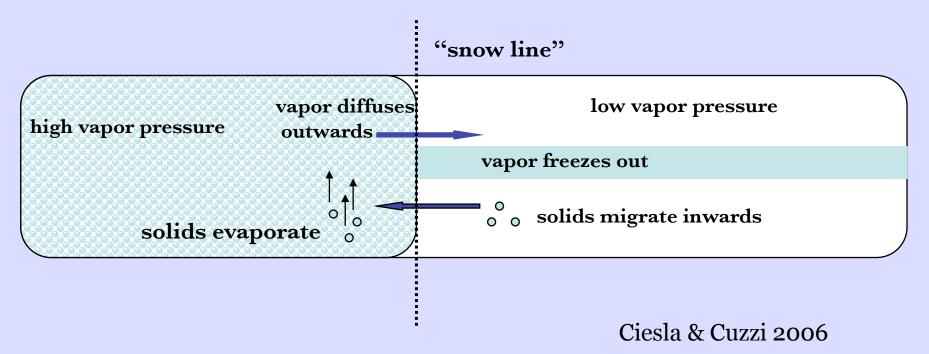
Delivery by comets or asteroids?



Drake & Righter 2002

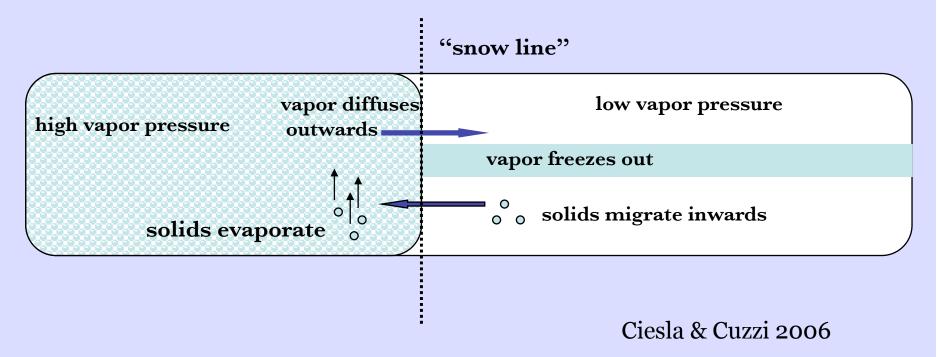
#### Water vapor as a tracer of disk evolution

Water vapor contents may be set by a balance between vapor diffusion and icy planetesimal migration



#### Water vapor as a tracer of disk evolution

Water vapor contents may be set by a balance between vapor diffusion and icy planetesimal migration



Inner disk water vapor content may reflect disk evolution

## Caveats for interpretation of emission spectra

To what extent do abundances depend on disk chemistry vs. disk transport?

Example: High UV fluxes may cause photodissociation of water at the disk surface

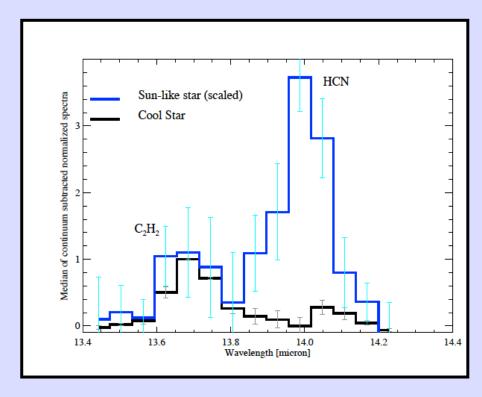
What are the line excitation mechanisms?

Example: Excitation models affect derived column densities

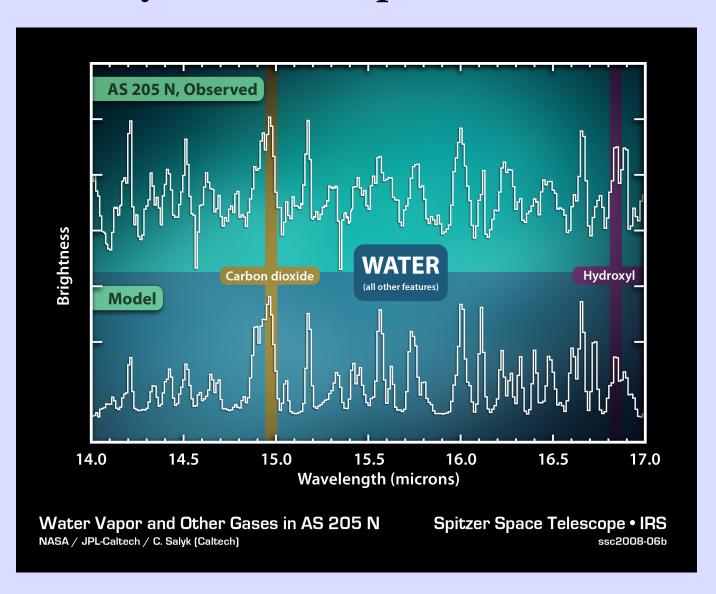
How are bulk disk properties expressed at the disk surface? Example: Dust settling and growth may enhance molecular emission

## Caveats for interpretation of emission spectra

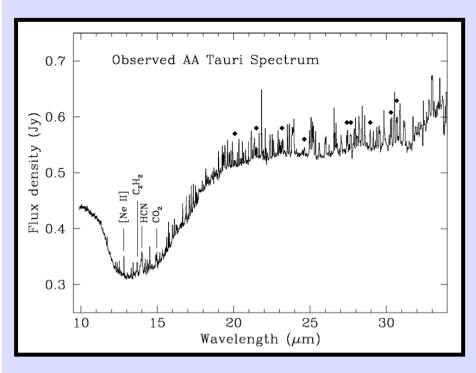
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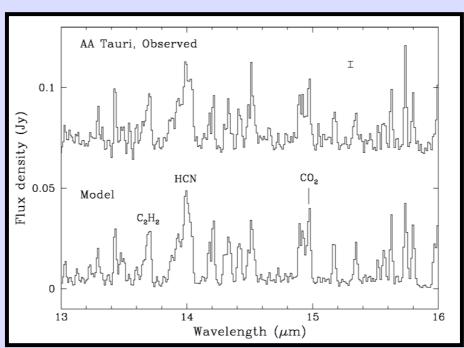


## Discovery of water vapor in inner disks



## Water (and other molecules) seen in many disks, with high enough S/N





Carr & Najita 2008

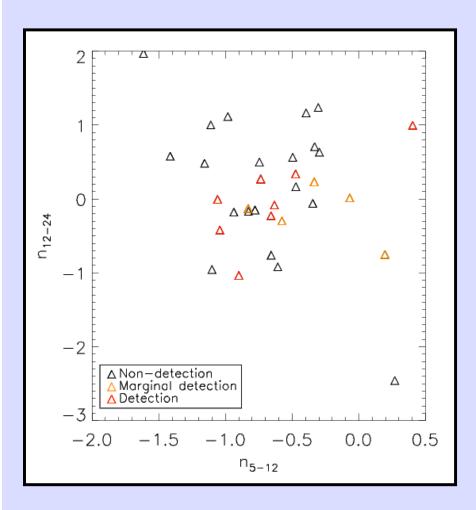
#### Molecular detections in c2d sources

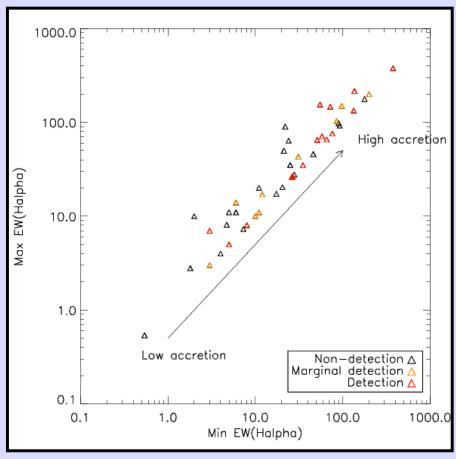
Water vapor detected in 30-50% of cTTSs

CO2 detected in many sources, with a variety of CO2/H2O ratios (including sources with CO2 but no H2O)

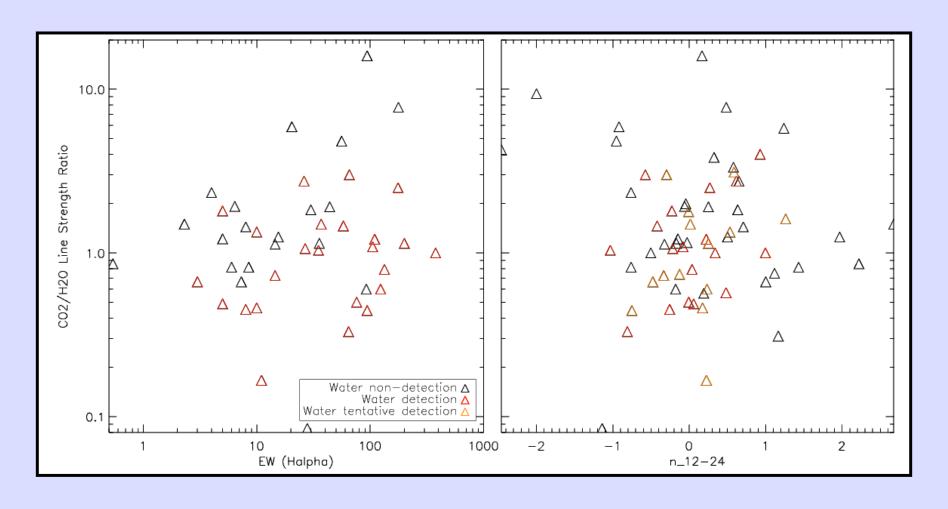
Many t entative detections of HCN and OH

#### Water detections in c2d sources

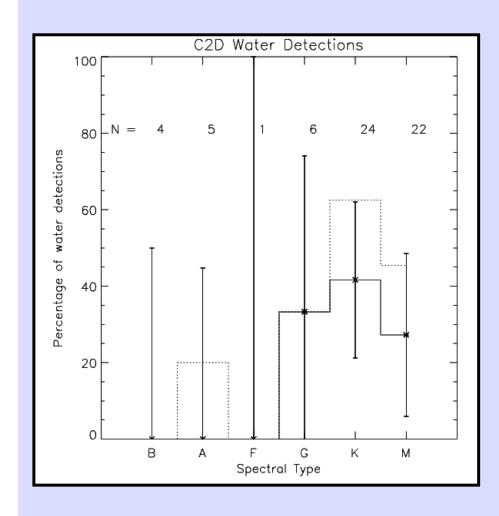


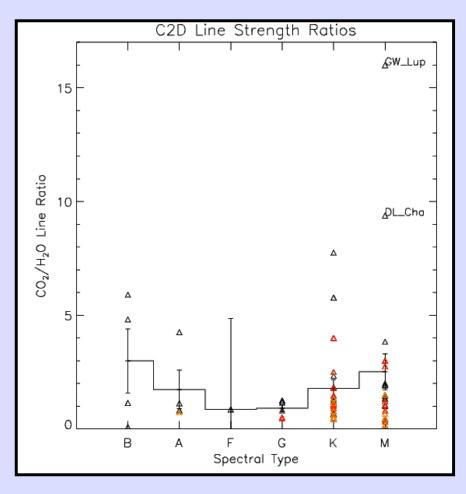


#### CO2 detections in c2d sources

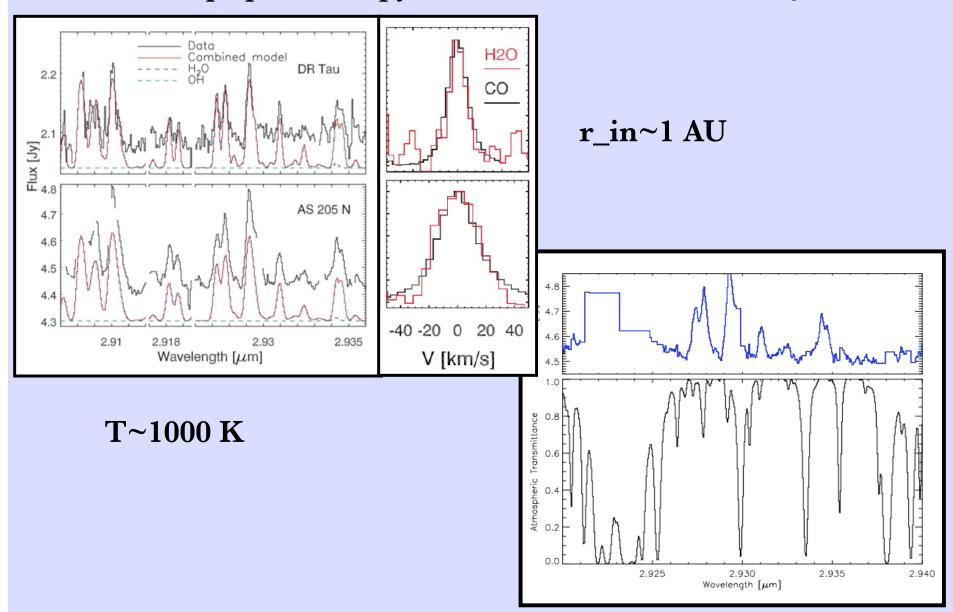


### **Spectral Type Dependencies**

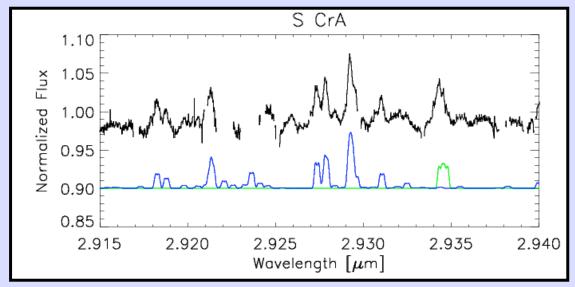


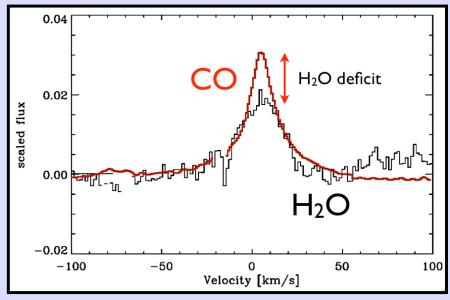


#### Follow-up spectroscopy with Keck-NIRSPEC (R~25,000)

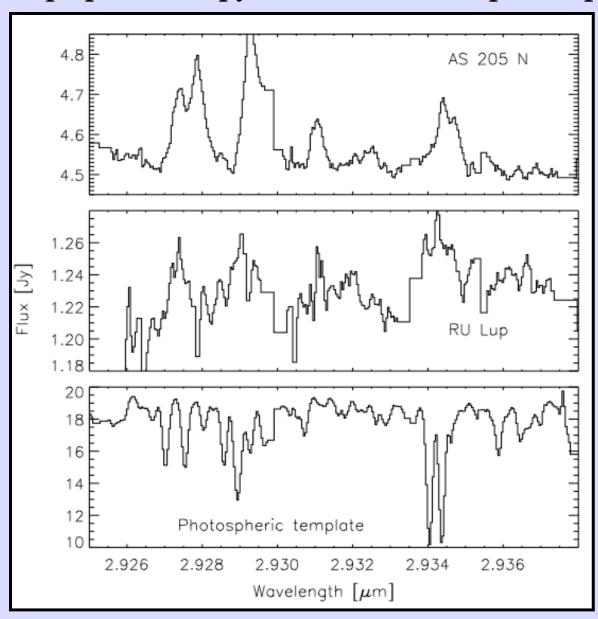


#### Follow-up spectroscopy with VLT-CRIRES (R~100,000)

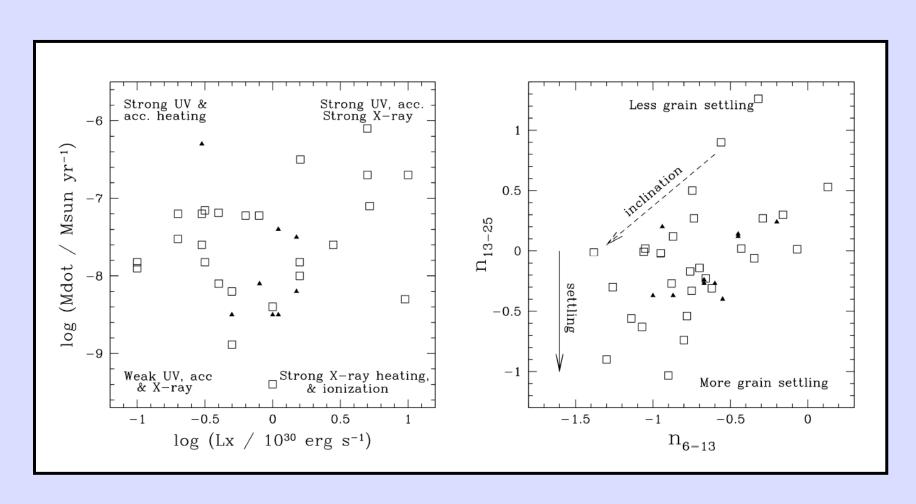




#### Follow-up spectroscopy: Effect of stellar photospheres

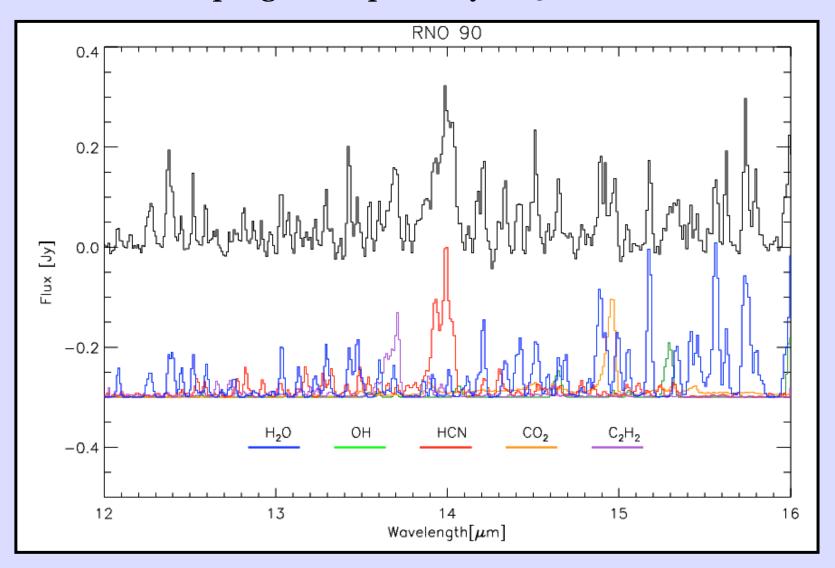


## Work in progress: Spitzer cycle 5 observations Survey of 41 disks in nearby star-forming regions



PID#50641, J. Carr, PI

#### Work in progress: Spitzer cycle 5 observations



#### Conclusions

Water vapor and other molecules common in IRS spectra. Emission lines tentatively more common for lower-mass stars.

High-resolution follow-up observations (NIRSPEC/CRIRES) suggest emission from small radii (r< a few AU).

A great variety of molecular ratios is seen, suggesting significant disk diversity. No trends observed just yet.

Cycle 5 high S/N spectra of 41 disks currently being processed. Most sources show rich molecular spectra; at least 80% show water emission.

Table 1: Fits to NIRSPEC data

	DR Tau	AS 205A
$CO v_{in}, v_{3\sigma} [km s^{-1}]$	27, 32	35, 40
CO $r_{\rm in}, r_{3\sigma}$ [AU]	0.8,  0.6	0.5,  0.4
$H_2O v_{\rm in}, v_{3\sigma} [{\rm km  s^{-1}}]$	24, 28	36, 42
$H_2O r_{in}, r_{3\sigma} [AU]$	1.0, 0.7	0.4,  0.3
OH $v_{\rm in}, v_{3\sigma} \ [{\rm km  s^{-1}}]$	28, 32	34, 39
OH $r_{\rm in}, r_{3\sigma}$ [AU]	0.7,  0.5	0.5,  0.4
r[AU] <sup>a</sup>	3	3 в
$\sigma~[{\rm km~s^{-1}}]$ a	2	2
$M_{\star}~[M_{\odot}]$ a	0.76	1.2
$i [^{\circ}]^{\mathrm{a}}$	67	47
$\Omega_{\rm CO, H_2O, OH} [\rm sr]$	$2 \times 10^{-16}$	$6 \times 10^{-16}$
$T_{\rm CO,H_2O,OH}$ [K]	1000	1000
$N_{\mathrm{CO}}$ [cm <sup>-2</sup> ]	$7  imes 10^{18}$	$6 \times 10^{18}$
$N_{ m H_{2}O}~{ m [cm^{-2}]}$	$8 \times 10^{17}$	$6 \times 10^{17}$
$N_{\rm OH}$ [cm <sup>-2</sup> ]	$2 \times 10^{17}$	$2 \times 10^{17}$

#### Possible solutions

D/H ratios observed in comets do not represent the bulk of the comet (Podolak, Mekler & Prialnik 2002)

D/H ratios observed in comets are probing the wrong reservoir

Adsorption of water onto grains followed by 'wet' accretion (Drake 2005)

Evolution of D/H ratio in Earth's oceans (Genda & Ikoma 2008)

### Problems with water delivery

