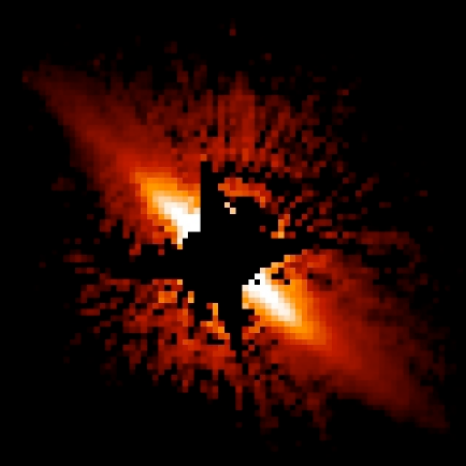
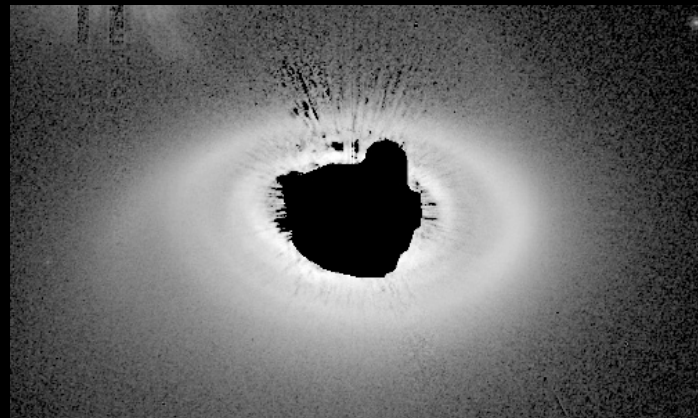
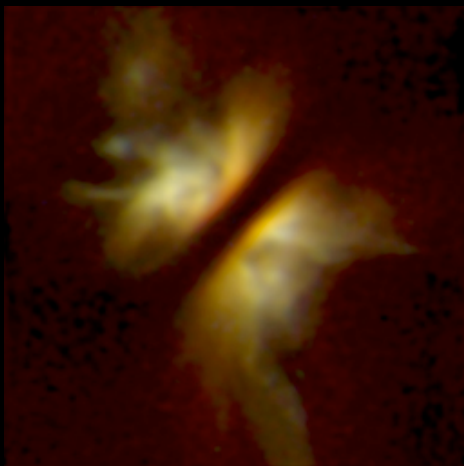
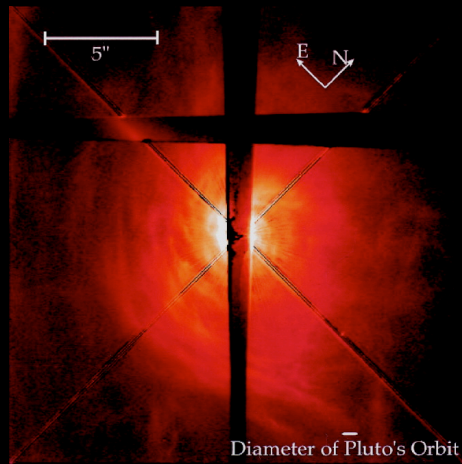
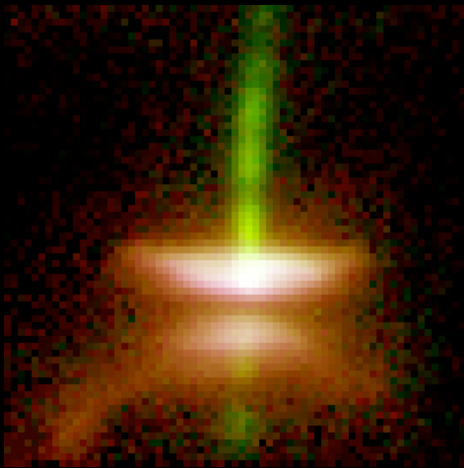


Gas in Protoplanetary Disks

Aki Roberge (NASA GSFC)

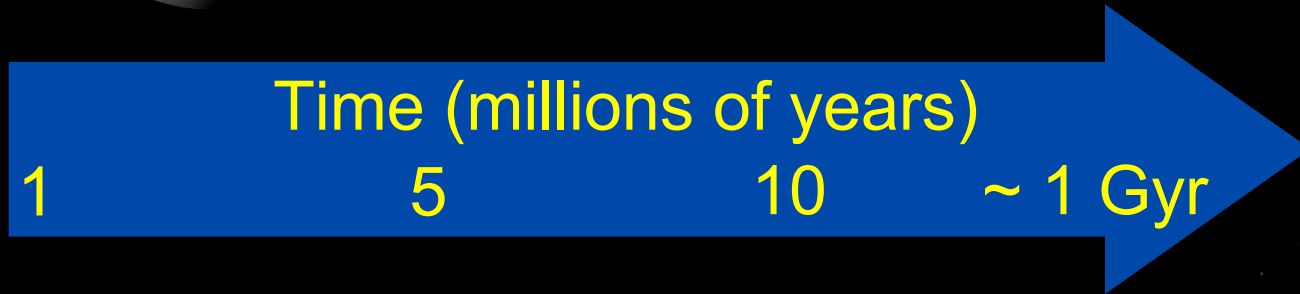


Gas in Protoplanetary Disks

Dr. Aki Roberge

Theoretical & Observational Timelines

Theory



Observation

Disk Evolution

$10 - 100 M_{\text{Jupiter}}$ $\text{few } M_{\text{Lunar}} \text{ (dust)}$

Total Mass

$10 - 100 M_{\text{Jupiter}}$?

Gas Mass

Time (millions of years)

1 5 10 ~ 1 Gyr



Observation

Giant Planet Formation & Migration



- Gas lifetime limits timescale for giant planet formation
- Causes migration (leading to hot Jupiters)

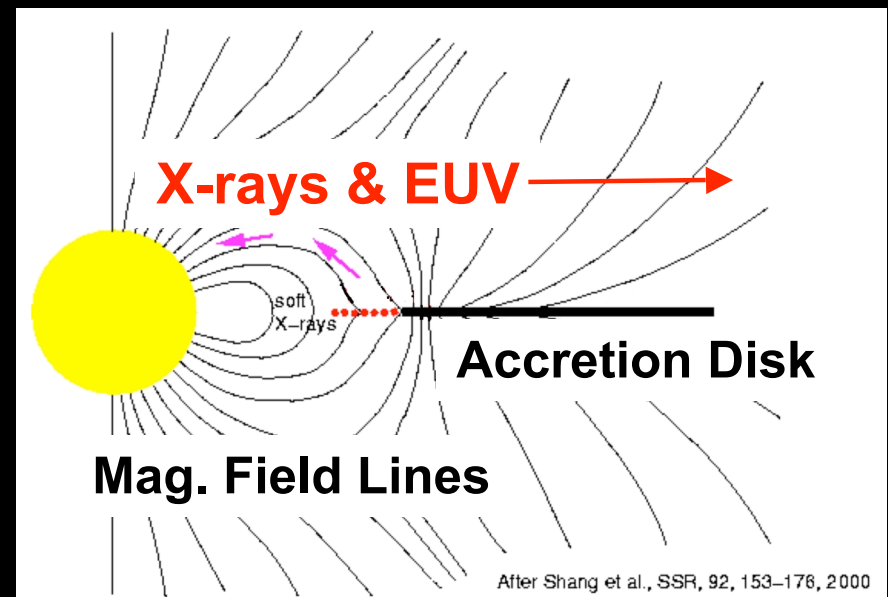


Measuring Gas Masses

- Two ways of estimating gas mass :
 - 1) Dust mass \times interstellar gas:dust ratio (100:1)
 - 2) CO mass \times interstellar H₂:CO ratio
- But disk gas:dust \neq interstellar gas:dust
 - Disk gas:dust ratio : 100 (primordial) \rightarrow \sim 0 (debris)
- Problems with CO technique :
 - 1) Sub-mm ¹²CO emission lines often saturate
 - 2) H₂:CO ratio sensitive to UV, freeze-out of CO

Mid-IR Gas Emission

- Few mid-IR H₂ detections
 - Probably due to low line-continuum contrast
 - Only hot ($T > 500$ K) gas lines
- Ne II emission @ 12.8 μm
 - 20% of CTTS w/ spectra from c2d Legacy program (Lahuis et al. 2007)



Upcoming Primordial Gas Studies

- Large dispersion in estimates of gas lifetimes
 - Between ~ 10 Myr and < 1 Myr (Hillenbrand 2005)
- Herschel Space Observatory
 - Far-IR to sub-mm imaging and spectroscopy



<http://herschel.esac.esa.int/>

Herschel Open Time Key Project

- “Gas in Protoplanetary Systems” (GASPS)

PI: W. Dent (UKATC)

– 400 hours for about 200 late B to M stars,

ages = 1 – 30 Myr

- [C II], [O I], H₂O
- Disk gas lifetimes,
total disk masses

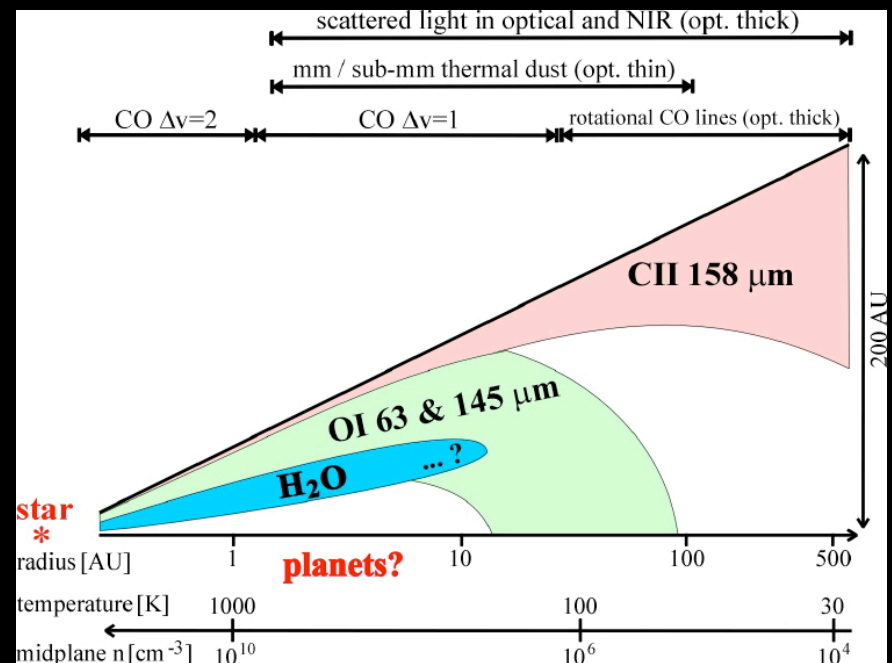
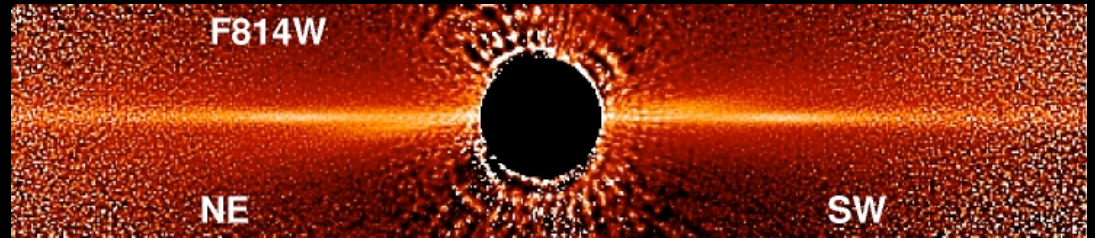


Image credit : I. Kamp

Secondary Gas in Debris Disks

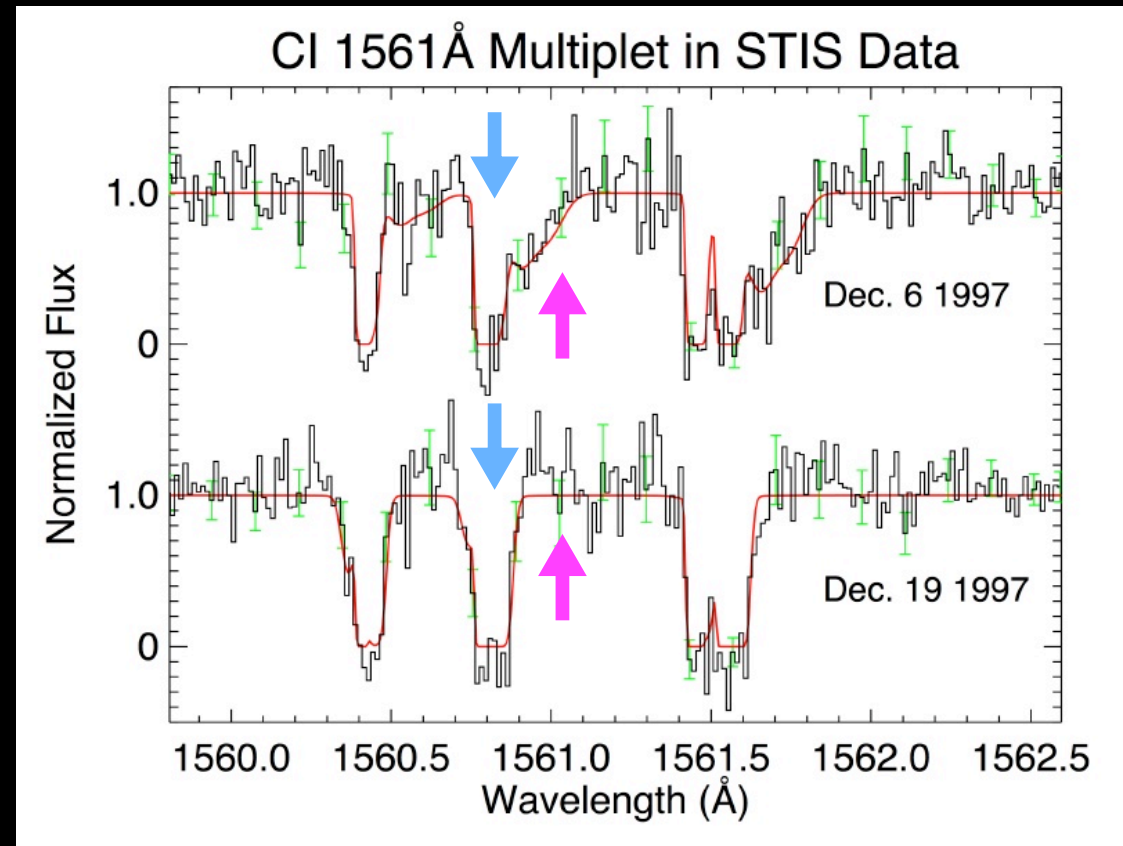
- Debris gas from planetesimals
 - Low gas:dust ratios
 - Mostly atomic gas



Golimowski et al. (2006)

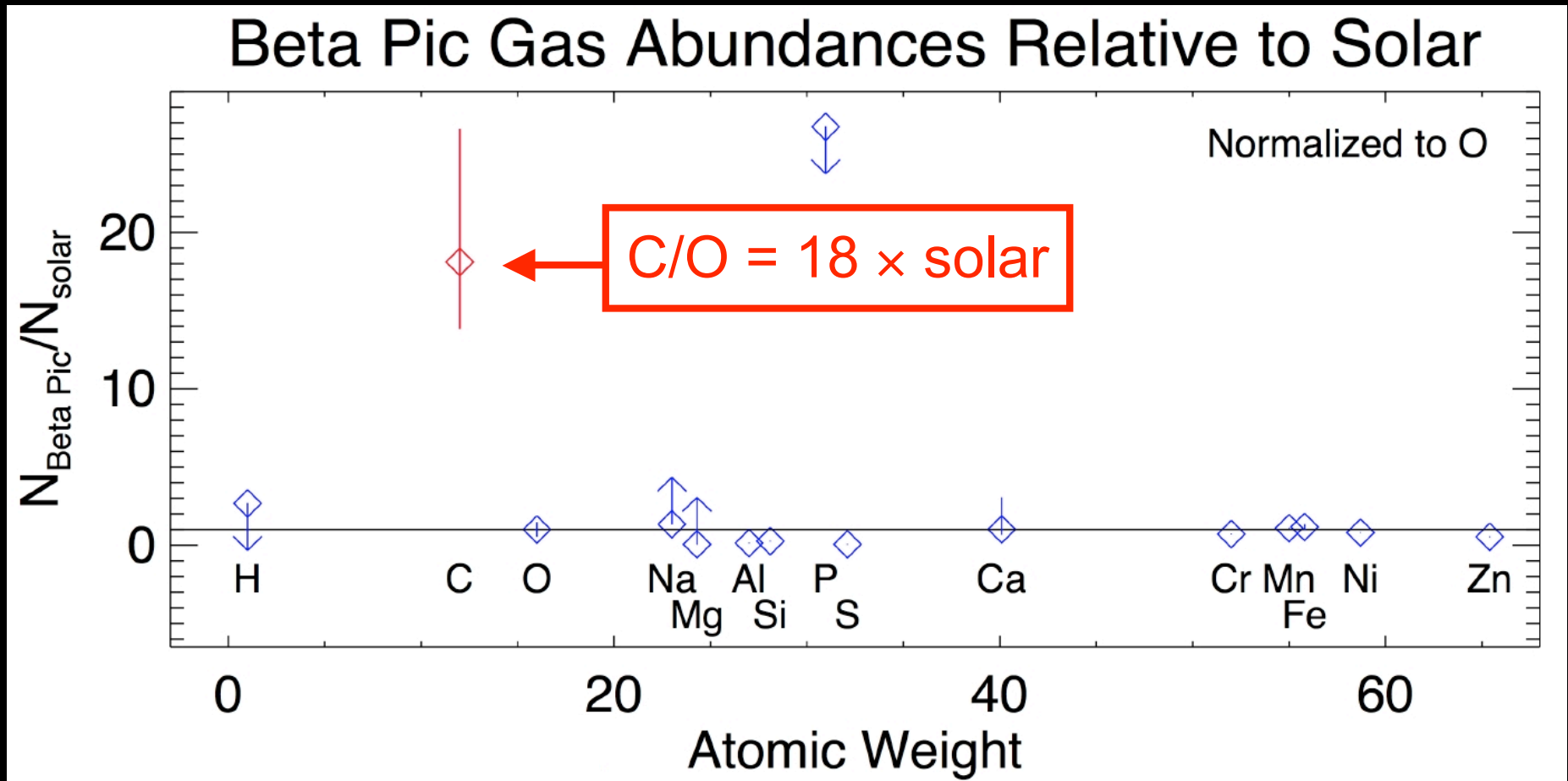
Secondary Gas in Debris Disks

- Debris gas from planetesimals
 - Low gas:dust ratios
 - Mostly atomic gas
- Narrow unvarying features at $v = v_{\star}$: stable gas
- Variable redshifted features : star-grazing planetesimals (FEBs)



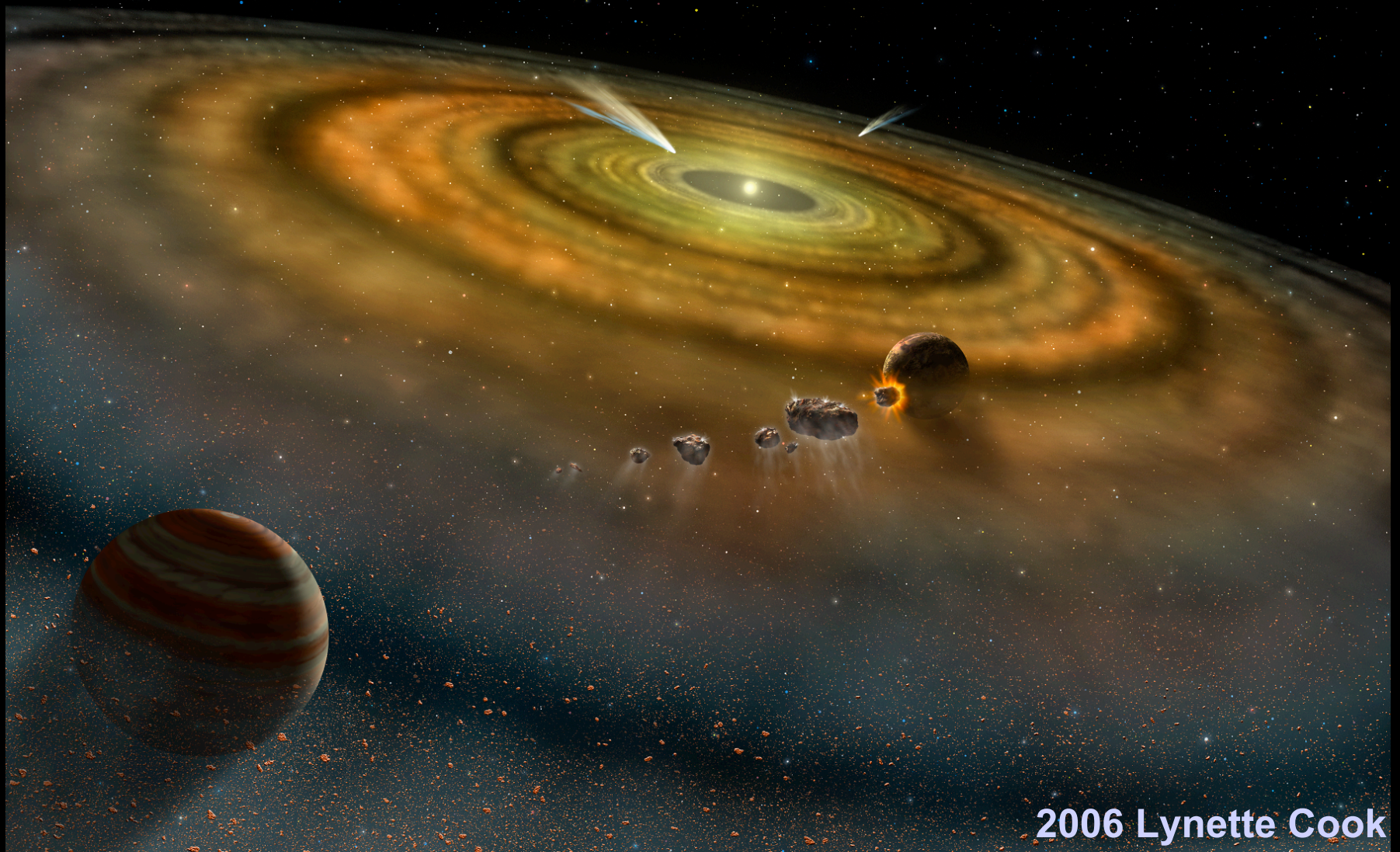
Roberge et al. (2000)

Beta Pic Gas Composition



Roberge et al. (2006)

C-rich AGB stars: $C/O < 1.2$
(Mattsson et al. 2007)



2006 Lynette Cook